

Effective LTB models

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Main Question

How do quantum gravity effects influence black holes?

Extensive literature with contributions from many researchers:

[Ashtekar, Bojowald, Modesto, Cartin, Khanna, Boehmer, Vandersloot, Chiou, Campiglia, Gambini, Pullin, Sabharwal, Brannlund, Kloster, De Benedictis, Olmedo, Dadhich, Joe, Singh, Haggard, Rovelli, Vidotto, Corichi, Saini, Cortez, Cuervo, Morales-Técotl, Ruelas, Pawlowski, Bianchi, Giesel, Christodoulo, D'Ambrosio, Alesci, Bahrami, Pranzetti, Husain, Kelly, Santacruz, Wilson-Ewing, Lewandowski, Zhang, Ma, Song, Bodendorfer, Mele, Münch, Navascués, Mena Marugán, García-Quismondo, Perez, Speziale, Viollet, Han, Liu, Alonso-Bardaji, Brizuela, Vera,...]

- Eternal BH models: Based on results of LQC study quantum corrections Reviews: [Gambini, Olmedo, Pullin '22], [Ashtekar, Olmedo, Singh '23]
- Dynamical process of gravitational collapse?
- Typical playground: spherical symmetric model with dust (perfect fluid, no pressure),
 i.e. in classical GR Lemaître-Tolman-Bondi spacetimes
- Effective description: classical model with correction functions

Introduction



- Most models describe Oppenheimer-Snyder scenario: homogeneous dust ball embeddded in vacuum
- Use classical junction conditions to glue interior Friedmann model to exterior
- Question: Is there a discontinuity in the gravitational field after the bounce? [Achour, Brahma, Uzan '20]
- Can we have decoupled equations of motion as in classical case?

We want to contribute to this discussion by

- Construction of effective LTB models from underlying effective spherical symmetric model as a 1+1 field theory under certain assumptions
- Start with general ansatz for effective model (not only motivated from $\overline{\mu}$ -scheme and its reduced quantization [Chiou, Ni, Tang '12], [Gambini, Olmedo, Pullin '20])
- In this setup we can have arbitrary dust mass profiles
- Can consider different coordinates (aerial gauge) due to underlying spherical symmetric model to relate to other models

Plan of the talk



- 1. Classical LTB in the canonical framework with connection variables
- 2. Constructing effective LTB models
 - Analyze stability of LTB condition in effective dynamics
- 3. Concrete model: Adapt to improved LQC dynamics
 - Analyze solution in marginal bound case
 - Extended mimetic gravity is underlying Lagrangian
- 4. Polymerized vacuum solution and comparison to other models
- 5. Summary and Outlook

Classical LTB model

classical LTB sector in spherical sym. spacetimes



Work with real Ashtekar-Barbero variables, Hamiltonian has standard form

$$H = \int \mathrm{d}x \left(NC + N^x C_x + \lambda G \right)$$

- Impose spherical symmetry on triad and connection (A,E) [Bojowald Kastrup '00], [Bojowald, Swiderski '06]
- Spherical symmetric metric has form (with $(E^{\phi})^2 = (E^1)^2 + (E^2)^2$)

$$ds^{2} = -N(x,t)^{2}dt^{2} + \frac{(E^{\phi})^{2}}{|E^{x}|}(dx + N^{x}dt)^{2} + |E^{x}|d\Omega^{2}.$$

- Can fix Gauß constraint to get rid off remaining gauge freedom
- Reduced phase space can be coordinatized by [Modesto '04]

$$\{K_x(x), E^x(y)\} = G\delta(x, y)$$
 $\{K_\phi(x), E^\phi(y)\} = G\delta(x, y)$

ullet Add dust to the gravitational system: (T,P_T)

classical LTB sector in spherical sym. spacetimes



Lemaitre-Tolman-Bondi sector

The LTB solution in these variables is given by [Lemaître '33], [Tolman '34], [Bondi '47]

$$ds^{2} = -dt^{2} + \frac{((E^{x})')^{2}}{4|E^{x}|(1+\mathcal{E}(x))}dx^{2} + |E^{x}| d\Omega^{2},$$

with the dynamical equation for each shell

$$\partial_t E^x = \pm 2\sqrt{E^x} \sqrt{\mathcal{E}(x) + \frac{\mathcal{F}(x)}{(E^x)^2}}.$$

Comparing LTB to general spherical symmetric metric, we need

$$N = 1$$
 $N^x = 0$ $G_x(x) = \frac{E^{x'}}{2E^{\phi}}(x) - \sqrt{1 + \mathcal{E}(x)} = 0$

Gauge Fixings

The LTB sector can be reached by the two gauge fixings

$$\left(C \longrightarrow G_T = T(x) - t\right), \qquad \left(C_x \longrightarrow G_x = \frac{E^{x'}}{2E^{\phi}}(x) - \sqrt{1 + \mathcal{E}(x)}\right)$$

Note: In marginally bound case ($\mathcal{E}=0$), in dust time gauge $\{G_x(x), C_x^{\text{tot}}(y)\}\approx 0$

Effective LTB models

Effective primary Hamiltonian



Start with partially gauge fixed effective system (dust time gauge):

Effective primary Hamiltonian

Consider effective model with temporal gauge fixed primary Hamiltonian

$$H_P^{\Delta}[N^x] = \int dx \, (C^{\Delta} + N^x C_x)(x) \,, \qquad C_x = \frac{1}{G} (E^{\phi} K_{\phi}' - K_x(E^x)')$$

and the polymerized gravitational contribution of the scalar constraint

$$C^{\Delta}(x) = \frac{E^{\phi}}{2G\sqrt{E^{x}}} \left[-(1+f)E^{x} \left(\frac{4K_{x}K_{\phi}}{E^{\phi}} + \frac{K_{\phi}^{2}}{E^{x}} \right) + h_{1} \left(\left(\frac{E^{x'}}{2E^{\phi}} \right)^{2} - 1 \right) + 2\frac{E^{x}}{E^{\phi}} h_{2} \left(\frac{E^{x'}}{2E^{\phi}} \right)' \right].$$

The polymerization functions have classical limit

$$h_1(E^x) \to 1$$
 $h_2(E^x) \to 1$ $f(K_x/E^{\phi}, K_{\phi}, E^x) \to 0$

Note: density weight unchanged since combination K_x/E^ϕ in f, thus

$$\left\{ C^{\Delta}[N], C_x[N^x] \right\} = C^{\Delta}[N^x(\partial_x N)]$$

⇒ Investigate dynamically stable reductions to LTB sector



Conservation of C^{Δ} (see Lemma 1 in [Giesel, Liu, Rullit, Singh, SW '23)

We have [Tibrewala '12], [Alonso-Bardaji, Brizuela '21]

$$\left\{ H_P^{\Delta}[N^x = 0], C^{\Delta}(y) \right\} \Big|_{C_x = 0} = 0,$$

if there is **no polymerization** of K_x and

$$\frac{h_1 - 2E^x \partial_{E^x} h_2}{h_2} = \frac{-4E^x \partial_{E^x} f^{(2)} + \partial_{K_\phi} f^{(1)}}{2f^{(2)}}.$$

The polymerization functions are defined in this case as

$$C^{\Delta}(x) = -\frac{E^{\phi}}{2G\sqrt{E^x}} \left[\frac{4K_x f^{(2)}(K_{\phi}, E^x)}{E^{\phi}} + \frac{f^{(1)}(K_{\phi}, E^x)}{E^x} - h_1 \dots \right] (x)$$

Note: this is not equivalent with closure of constraint algebra since this should be analyzed in fully gauge unfixed system

Implementing polymerized LTB condition



• Introduce effective LTB condition (classical: $G_x^{\Delta} = \frac{E^{x'}}{2E^{\phi}} - \sqrt{1+\mathcal{E}}$)

$$G_x^{\Delta} = \frac{E^{x'}}{2E^{\phi}} - g_{\Delta}(K_x/E^{\phi}, K_{\phi}, E^x, [\partial_x^n(K, E)], \mathcal{E})$$

- Investigate stability of effective LTB condition under effective dynamics [Bojowald, Harada, Tibrewala '08], [Bojowald, Reyes, Tibrewala '09]
 - We call such LTB conditions compatible
- Our strategy: work on the level of equations of motion

Question

For which G_x^{Δ} do the four EOM of $\dot{K}_x, \dot{K}_\phi, \dot{E}^x, \dot{E}^\phi$ reduce to only **two** in the sector

$$N^x = 0, C_x = 0, G_x^{\Delta} = 0$$
 ?

First result: compatible LTB conditions are of the form

$$g_{\Delta} = g_{\Delta}^{(1)}(K_{\phi}, E^x, \mathcal{E}) + g_{\Delta}^{(2)}(\widetilde{K}_x = \frac{\partial_x K_{\phi}}{\partial_x E^x}, K_{\phi}, E^x)$$

ullet Contribution $g_{A}^{(1)}$ represents non-marginally and $g_{A}^{(2)}$ marginally bound case

Implementing polymerized LTB condition



Key results on the form of the polymerization functions and compatible LTB condition:

Key results

- 1. Additional condition for K_x polymerization
 - \Rightarrow Non-marginal: always no K_x polymerization allowed
 - \Rightarrow Marginal: if $g_{A}^{(2)}=g_{A}^{(2)}(K_{\phi},E^{x})$ no K_{x} polymerization allowed
- 2. Separating LTB function dependence $g_{\Delta}^{(1)} = \tilde{g}_{\Delta}(K_{\phi}, E^x) \sqrt{1 + \mathcal{E}}$

$$g_{\Delta}^{(1)} = \tilde{g}_{\Delta}(E^x)\sqrt{1+\mathcal{E}} \qquad 1 - \frac{2E^x\partial_{E^x}\tilde{g}_{\Delta}}{\tilde{g}_{\Delta}} = \frac{-4E^x\partial_{E^x}f^{(2)} + \partial_{K_{\phi}}f^{(1)}}{2f^{(2)}}$$

second condition allows the conservation of C^{Δ} when we further have

$$2E^x \partial_{E^x} \tilde{g}_{\Delta} = \left(1 - \frac{h_1 - 2E^x \partial_{E^x} h_2}{h_2}\right) \tilde{g}_{\Delta}.$$

3. Classical LTB condition restricts inverse triad and holonomy corrections to

$$\partial_{K_\phi} f^{(1)} = 2f^{(2)} + 4E^x \partial_{E^x} f^{(2)} \qquad \qquad h_1 = h_2 + 2E^x \partial_{E^x} h_2$$
 "conservation"



Consider system with compatible LTB condition and C^{Δ} conserved (Lemma 1):

Dynamical equations (see Corollary 4 in [Glesel, Liu, Rullit, Singh, SW '23])

The dynamics of such models decouple in radial x-direction

$$\partial_t E^x = 2\sqrt{E^x} f^{(2)}$$

$$\partial_t K_\phi = -\frac{1}{2\sqrt{E^x}} \left(f^{(1)} - \tilde{g}_\Delta^2 (1 + \mathcal{E})(2h_2 + 4E^x \partial_{E^x} h_2 - h_1) + h_1 \right)$$

- Equations applicable in marginally and non-marginally bound case
- Non-marginal: same result as implementing gauge fixing G_x^Δ and computing associated Dirac bracket
- Result supports assumption of decoupled shells in dust collapse models, e.g. [Kiefer, Schmitz '19], [Giesel, Li, Singh'21]
- Solution parametrized by energy $\mathcal{E}(x_0)$ and conserved quantity mass $M(x_0)$ of a shell at $x=x_0$



Matching LQC dynamics



General strategy

- Choose an effective LQC model as starting point
- Use Cor. 4 to identify effective spherically symmetric model and LTB condition

In this way we get

- Underlying spherical symmetric model, that has no areal gauge implemented yet
- Dynamically stable reduction to LTB sector through effective LTB condition
- Equations of motion are decoupled and coincide with chosen LQC model

Sometimes one can relate effective spherical symmetric model to an underlying covariant Lagrangian

- In our model this will be extended mimetic gravity in comoving gauge
- Redefinition of time dependence of mimetic field corresponds to coordinate transformations in the temporal coordinate

Matching LQC dynamics



Adapt decoupled effective LTB sector to improved LQC dynamics [Ashtekar, Pawlowski, Singh '06]

$$\partial_t v = 3v \frac{\sin(2\alpha b)}{2\alpha}, \quad \partial_t b = -\frac{1}{2} \left(\frac{\mathcal{E}(x)}{v^{\frac{2}{3}}} + \frac{3\sin^2(\alpha b)}{\alpha^2} \right) ,$$

where we defined $v=(E^x)^{3/2}, b=\frac{K_\phi}{\sqrt{E^x}}, \alpha=\beta\sqrt{\Delta}$.

Underlying spherical symmetric model

The corresponding gauge unfixed effective spherical symmetric model is [Tibrewala '12]

$$C^{\Delta} = -\frac{E^{\phi}\sqrt{E^{x}}}{2G} \left[\frac{3}{\alpha^{2}} \sin^{2}\left(\frac{\alpha K_{\phi}}{\sqrt{E^{x}}}\right) + \frac{(2E^{x}K_{x} - E^{\phi}K_{\phi})}{\alpha\sqrt{E^{x}}E^{\phi}} \sin\left(\frac{2\alpha K_{\phi}}{\sqrt{E^{x}}}\right) + \frac{1 - \left(\frac{E^{x'}}{2E^{\phi}}\right)^{2}}{E^{x}} - \frac{2}{E^{\phi}} \left(\frac{E^{x'}}{2E^{\phi}}\right)' \right].$$

No inverse triad corrections: compatible LTB condition is classical one

$$G_x^{\Delta} = G_x = \frac{E^{x'}}{2E^{\phi}} - \sqrt{1 + \mathcal{E}}$$

and C^{Δ} is conserved quantity.



We can write dynamical equation as modified Friedmann equation

$$\frac{\dot{R}^2}{R^2}(x) = \left(\frac{\kappa\rho}{6} + \frac{\mathcal{E}}{R^2}\right) \left(1 - \alpha^2 \left(\frac{\kappa\rho}{6} + \frac{\mathcal{E}}{R^2}\right)\right) (x),$$

where the metric has the form (we work in LTB coordinates)

$$ds^2 = -dt^2 + (\partial_x R)^2 dx^2 + R^2 d\Omega^2.$$

Marginally bound case

In marginally bound case $\mathcal{E} = 0$ the solution is

$$R(x,t) = \sqrt{E^x} = \left(\mathcal{F}(x)\left(\frac{9}{4}(\tilde{\beta}(x) - t)^2 + \alpha^2\right)\right)^{\frac{1}{3}}$$

for homogeneous dust already in [Giesel, Han, Li, Liu, Singh '22], [Fazzini, Rovelli, Soltani '23]

- In vacuum: time symmetry, and metric is stationary
- No shell crossing singularities for vacuum and OS collapse, but in general inhomogeneous case not true [Fazzini, Husain, Wilson-Ewing '23]
- Horizons can form when $M(x)=rac{\mathcal{F}}{2G}>M_c=rac{8lpha}{3\sqrt{3}G}$ [Kelly, Santacruz, Wilson-Ewing '20], [Giesel, Han, Li, Liu, Singh '22], [Lewandowski, Ma, Yang, Zhang '22]

Extended mimetic model



Underlying covariant Lagrangian

Primary Hamiltonian can be generated from 2d action [Achour, Lamy, Liu, Noui '18], [Han, Liu '22]

$$S_2 = \frac{1}{4G} \int_{\mathcal{M}_2} d^2x \, \det(e) e^{2\psi} \left\{ \mathcal{R} + L_{\phi}(X, Y) + \frac{\lambda}{2} \left(\phi_{,j} \phi^{,j} + 1 \right) \right\},$$

in comoving gauge $\phi(t,x)=t$, note $\det(e)=E^{\phi}\sqrt{E^x}$ coupling

- (smooth) mimetic field naturally defines foliation into spacelike surfaces defined by $\phi=const.$, w.l.o.g. $\phi(x,t)=\phi(t)$
- Higher derivative coupling in X, Y relates to extrinsic curvature

$$X = -\Box_h \phi + Y = \frac{\partial_t E^{\phi}}{E^{\phi}}, \qquad Y = -h^{ij} \partial_i \psi \partial_j \phi = \frac{\partial_t E^x}{2E^x} = \frac{\sin(2\alpha b)}{2\alpha},$$

Pendant of the Einstein equations of this model are

$$G^{\Delta}_{\mu\nu} := G_{\mu\nu} - T^{\phi}_{\mu\nu} = -\lambda \partial_{\mu}\phi \partial_{\nu}\phi, \quad \partial_{\mu}\phi \partial^{\mu}\phi = -1.$$

Polymerized vacuum solution



Polymerized vacuum

Specialize to matter profile $\mathcal{F}=R_s=\mathrm{const}$ [Giesel, Han, Li, Liu, Singh '23], [Fazzini, Rovelli, Soltani '23]

$$R(x,t) = \left(R_s\left(\frac{9}{4}z^2 + \alpha^2\right)\right)^{\frac{1}{3}}, \qquad z := x - t$$

metric clearly stationary

• Corresponds to $\lambda = C^{\Delta} = 0$ but curvature non-vanishing

$$\mathcal{R} = -\frac{96\alpha^2}{\left(4\alpha^2 + 9z^2\right)^2}$$

due to non trivial coupling of ϕ in $T^\phi_{\mu\nu}\Rightarrow {\rm QG}$ effects

- Everything bounded, at bounce z = 0 no shell crossing singularity
- (smooth) signature change of E^{ϕ} at the bounce

$$E^{\phi}(t,x) = \frac{1}{2}(E^x)' = R(z)\partial_z R(z)$$

allowed due to $\det(e)$ coupling in Lagrangian and consistent with degeneracy of metric

Comparison to other models



Underlying covariant model allows to perform coordinate transformations in t and x

- ⇒ Relate to other models in the literature and discuss implications (shocks?)
 - Polymerized vacuum:
 - \circ Transform to Schwarzschild-like coordinates $(t,x) \to (\tau,z)$

$$ds^{2} = -A(r)d\tau^{2} + \frac{1}{A(r)}dr^{2} + r^{2}d\Omega^{2} \qquad A(r) = 1 - \frac{2Gm_{s}}{r} \left(1 - \frac{\alpha^{2}}{r^{2}} \frac{2Gm_{s}}{r}\right)$$

same solution as [Kelly,Santacruz,Wilson-Ewing '20],[Parvizi,Pawlowski,Tavakoli,Lewandowski '21],[Lewandowski,Ma,Yang,Zhang '22]

- \circ transformation only well defined for monotonic R(z) (not at bounce)
- Areal gauge:
 - \circ Gauge fix C_x with $G^{
 m ar}=E^x-r^2$, shift vector $N^r=-rac{r}{2lpha}{
 m sin}\Big(rac{2lpha}{r}K_\phi\Big)$
 - ⇒ We can exactly reproduce model in [Husain, Kelly, Santacruz, Wilson-Ewing '22]
 - We can transform our solution in LTB coordinates to Gullstrand-Painlevé

$$N^r = -\partial_t R(t,x) = -\mathrm{sign}(\partial_t R(t,x)) \sqrt{\frac{2GM(x)}{r}(...)}$$

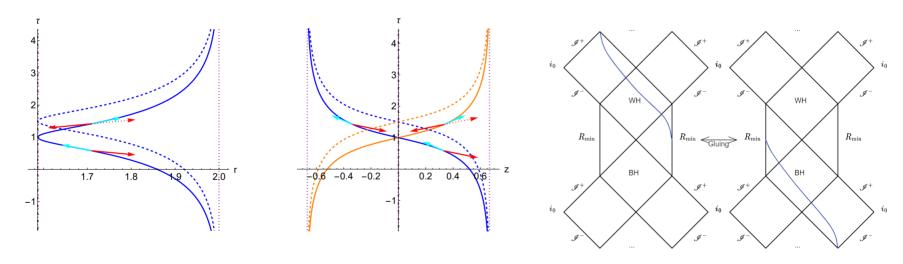
smooth signature change at bounce

Not observed when working directly in radial coordinates in vacuum case
 ⇒ gives rise to discontinuity/shocks in OS collapse

Comparison to other models



- Different global structure of spacetime can be seen in x = const geodesics
- These are also world lines of clock field ϕ , they intersect in (τ, r) after bounce \Rightarrow discontinuity in clock field [Fazzini, Rovelli, Soltani '23] violates smoothness of mimetic field
- To have same global structure in r coordinates need to consider gluing of two patches with different orientation [Münch '21]



 OS collapse: Gluing with effective junction conditions does not allow shocks without violating smoothness of mimetic field

Inhomogeneous dust collapse



- Consider inhomogeneous dust profile on initial slice
- In OS collapse everything bounded
 ⇒now shell crossing singularity[Fazzini,Husain,Wilson-Ewing '23]
- Prob. weak singularity: geodesics can pass through
- Separates spacetime into regions with different orientation
- ⇒ Needs further investigation, work in progress

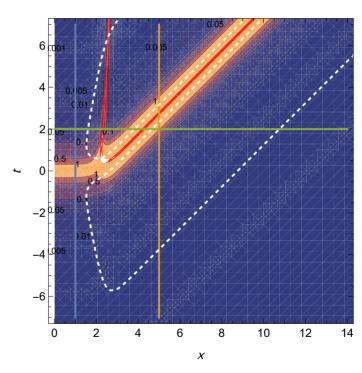


Figure: Kretschmann scalar

Conclusion



Summary

- Our framework allows construction of effective LTB models with holonomy and inverse triad corrections under certain assumptions (no polymerization of diffeo)
- Certain class of effective LTB models has decoupled dynamics
- LQC model as starting point: field theoretic model for inhomogeneous dust collapses
- Underlying mimetic model provides all coordinate transformations

Future work

- study further phenomenological properties like BH evaporation
- Extend analysis to LQC models with asymmetric bounce (work in progress)
- study polymerized vacua for more general polymerizations (work in progress)

Thank you for your attention!